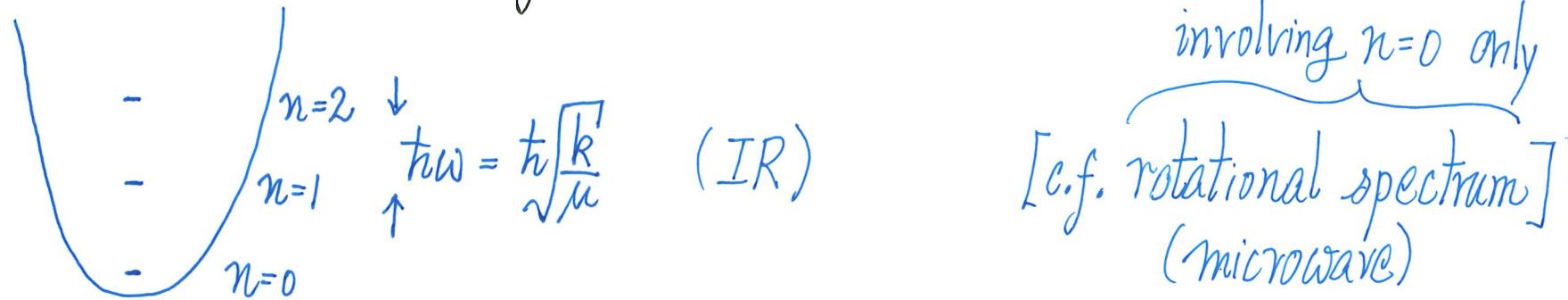


# I. Transitions involving different vibrational levels



Q: What will happen when molecules interact with IR light?

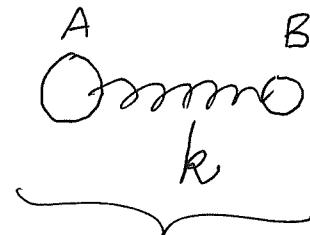
Q: Any condition on the molecular property for absorption?

- Depending on resolution, spectrum shows different details
- $n=0 \rightarrow n=1$  oscillator state typical [room temp  $\sim \frac{1}{40}$  eV]

under what condition would this happen?

How to use spectrum to extract  $k$  ("spring constant")?

# Diatomc Molecules



$$\omega = \sqrt{\frac{k}{\mu}}, \quad f = \frac{\omega}{2\pi}$$

only one normal mode [not counting freely moving CM]

## Fundamental Vibrational Frequencies and Effective Force Constants for Some Diatomic Molecules

Molecule	Frequency (Hz), $n = 0$ to $n = 1$	Force Constant (N/m)
diatomic molecules	HF	970
	HCl	480
	HBr	410
	HI	320
	CO	1860
	NO	1530

From G. M. Barrows, *The Structure of Molecules*, New York, W. A. Benjamin, 1963.

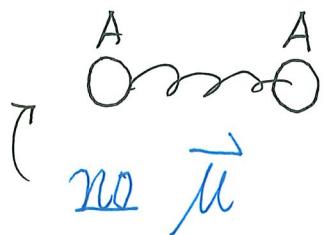
Units:  $5 \times 10^{13} \text{ Hz} = 50 \text{ THz}$ ;  $0.207 \text{ eV}$ ;  $\sim 6000 \text{ nm}$  (wavelength);  $1668 \text{ cm}^{-1}$  (wave number)  
 $6 \mu\text{m}$

- Why do we show  $\nu(\omega)$  for AB molecules with A  $\neq$  B?

good

Electric dipole mechanism,

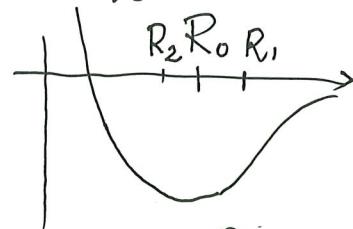
molecule should carry an electric dipole moment



"~~no~~" due to electrons and electrons  
will not bias any side

good

Is that all?



$$\mu = qR_0 \text{ (constant)}$$



$$\mu = qR_1$$



$$\mu = qR_2$$

Need this to have  
vibrational spectrum!

$\mu$  changes as  
atoms vibrate!

Aside: Transitions involving  $n' \leftrightarrow n$  involve [e.g.  $n=0, n'=1$ ]

$$\int \underbrace{\phi_{n'}^*(r-R_0)}_{\text{oscillator state } n'} \underbrace{[-\vec{\mu} \cdot \vec{E}_0]}_{\text{oscillator state } n} \underbrace{\phi_n(r-R_0) dr}_{\text{related to } \Delta l \text{ (after taking care of } \Delta n)} \cdot [\text{integral for } \Theta, \Phi]$$

if  $\vec{\mu}$  = constant (permanent dipole),  $\int \phi_{n'}^* \phi_n dr = 0$  ( $n \neq n'$ ) oscillator physics

Need  $\vec{\mu} \approx \underbrace{\vec{\mu}(R=R_0)}_{\text{constant}} + \underbrace{\frac{d\vec{\mu}}{dR} \Big|_{R=R_0}}_{\text{electric dipole changes as } R \text{ vibrates about } R_0 \neq 0} \cdot (R-R_0)$

then  $\int \phi_{n'}^*(r-R_0) (R-R_0) \phi_n(r-R_0) dr$

$$\propto \int \phi_{n'}^*(x) \cdot x \cdot \phi_n(x) dx \neq 0 \text{ only for } \Delta n = \pm 1$$

Also Needs molecule to have varying  $\vec{\mu}$  as it vibrates

Two results follow : (a) Need  $\frac{d\vec{\mu}}{dR} \Big|_{R=R_0} \neq 0$  (43)

Changing electric dipole moment  
as molecule vibrates

Important for science of global warming

(b)  $\Delta n = \pm 1$  (44)

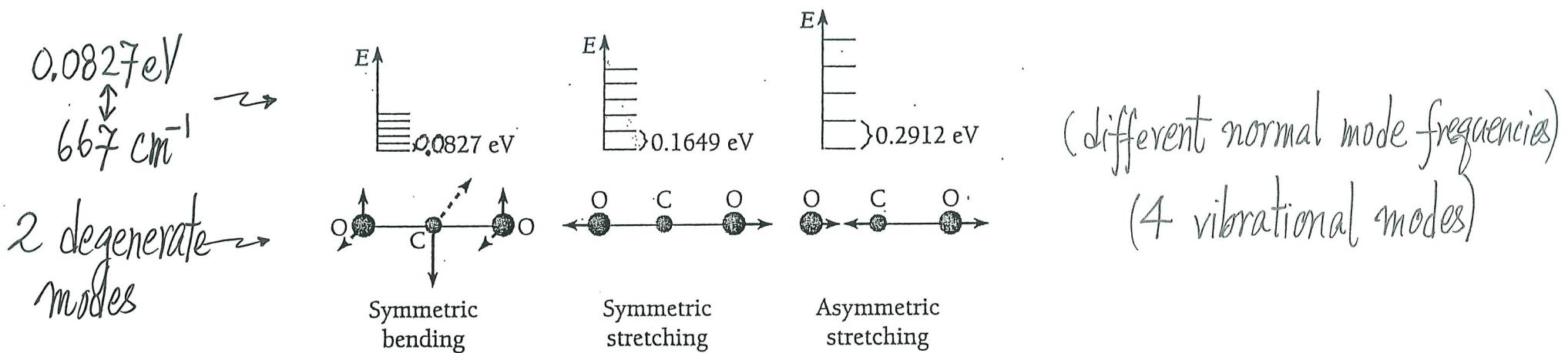
selection rule on transitions between  
vibrational levels

[there is an angular integral determining  $\Delta l$  after taking care of  $\Delta n$ ]

Results applicable to polyatomic molecules.



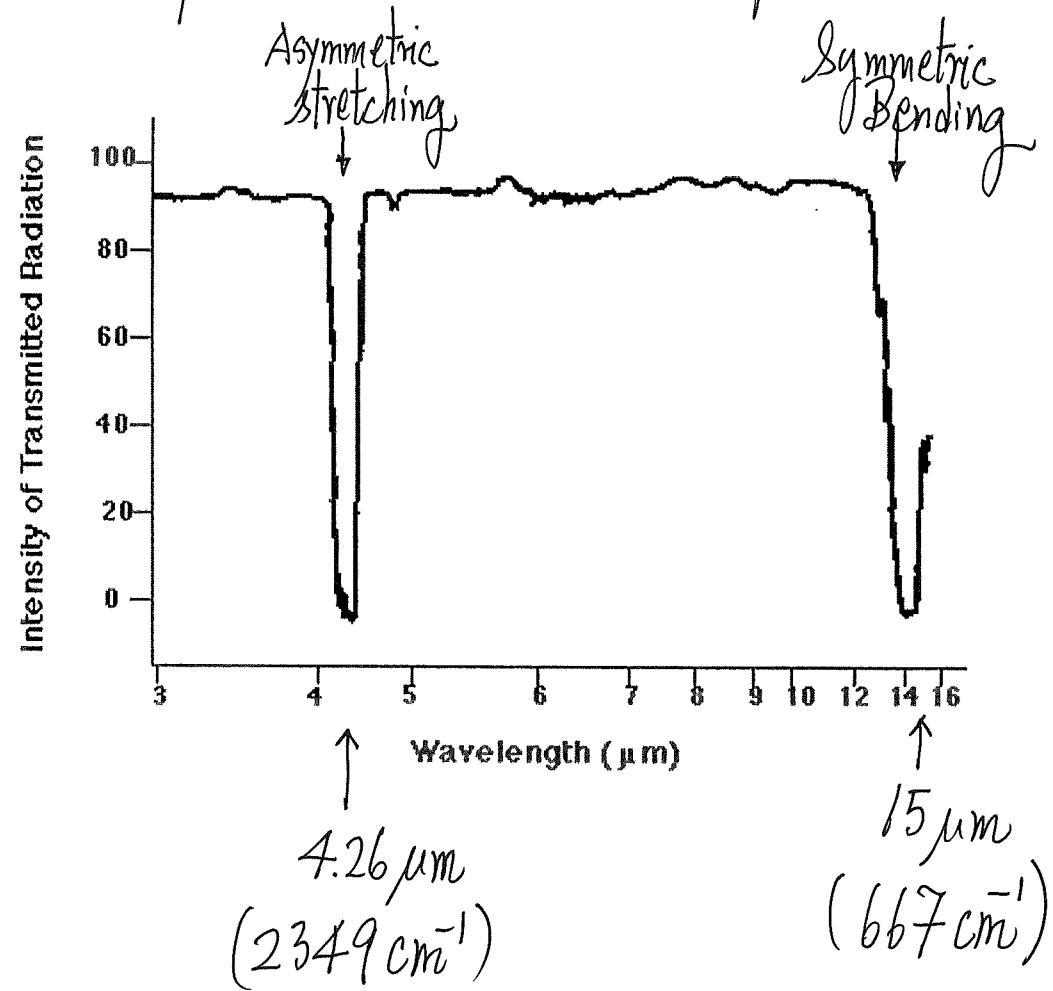
Normal modes, each with characteristic normal mode frequency  
 $\Rightarrow$  several independent oscillators



The normal modes of vibration of the  $\text{CO}_2$  molecule and the energy levels of each mode.  
The symmetric bending mode can occur in two perpendicular planes.

IR active	IR inactive	IR active
$667 \text{ cm}^{-1}$ $(15 \mu\text{m})$	[ $\bar{\mu}$ doesn't change] $1388 \text{ cm}^{-1}$ $(7.2 \mu\text{m})$	$2349 \text{ cm}^{-1}$ $(4.26 \mu\text{m})$

$\text{CO}_2$  spectrum in IR [absorption due to IR active modes]

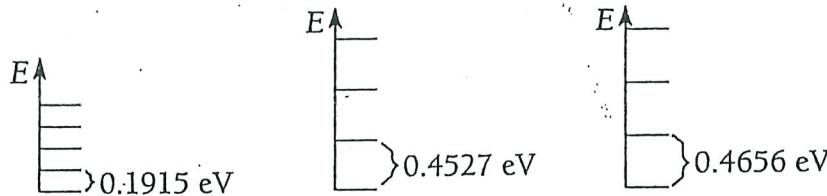


# H<sub>2</sub>O: Normal Modes (H<sub>2</sub>O is not a linear molecule)

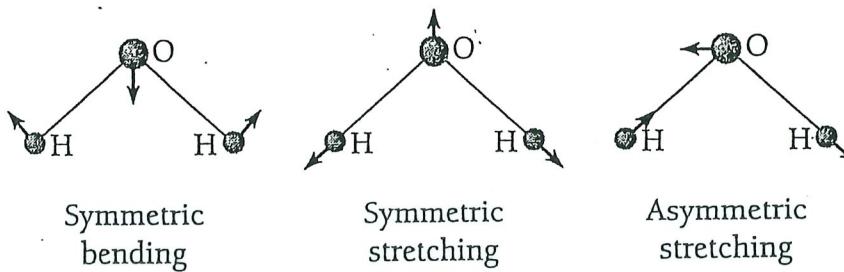
$$0.1915 \text{ eV}$$

$\Downarrow$

$$\sim 1540 \text{ cm}^{-1}$$

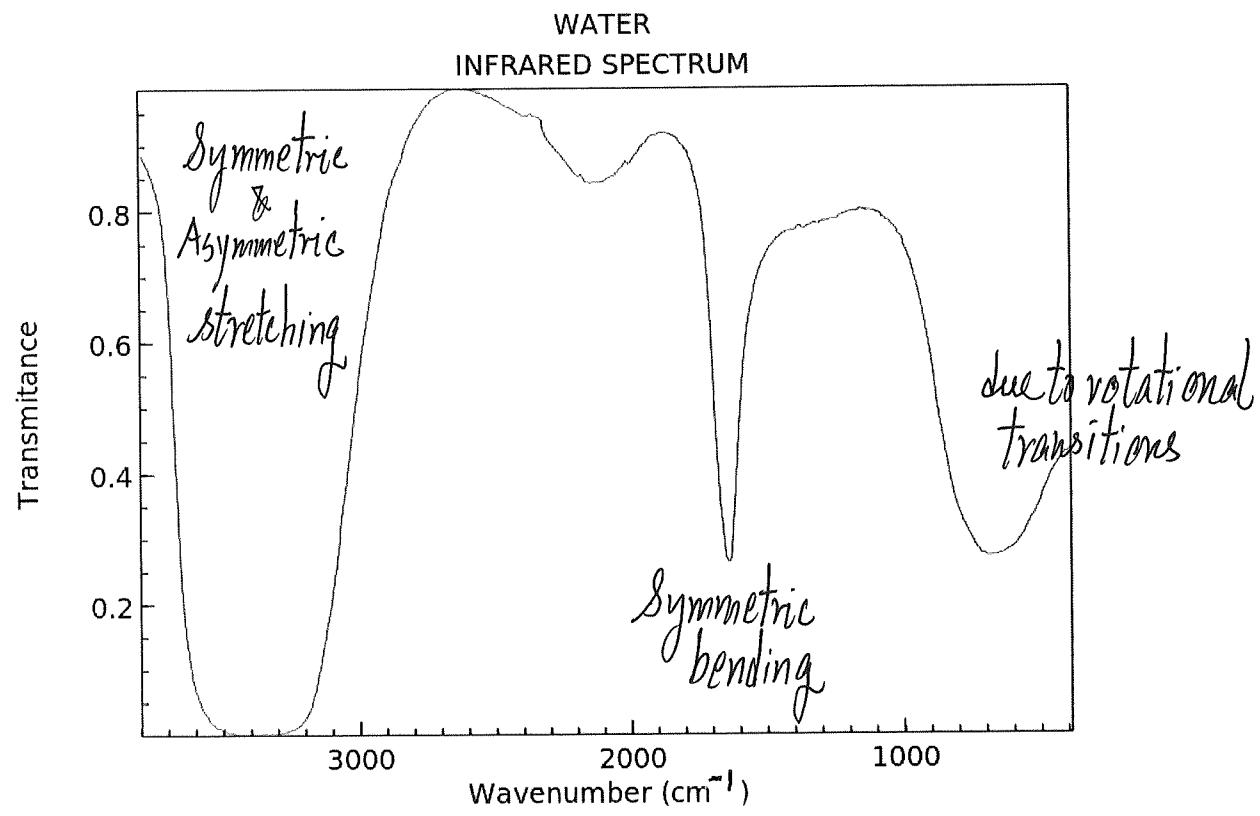


3 normal modes



↑ ALL are IR active ↑ ( $\vec{\mu}$  changes as molecule vibrates)

$$\begin{array}{ccc} \uparrow & \uparrow & \uparrow \\ 1595 \text{ cm}^{-1} & 3657 \text{ cm}^{-1} & 3756 \text{ cm}^{-1} \\ (6.27 \mu\text{m}) & (2.74 \mu\text{m}) & (2.66 \mu\text{m}) \end{array}$$

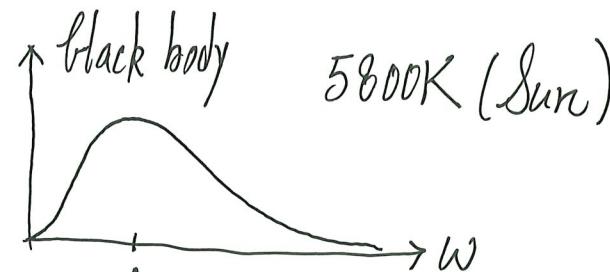


NIST Chemistry WebBook (<https://webbook.nist.gov/chemistry>)

$\text{CO}_2$  and  $\text{H}_2\text{O}$  (vapour) are Greenhouse Gases

Sun  $T \sim 5800\text{K}$

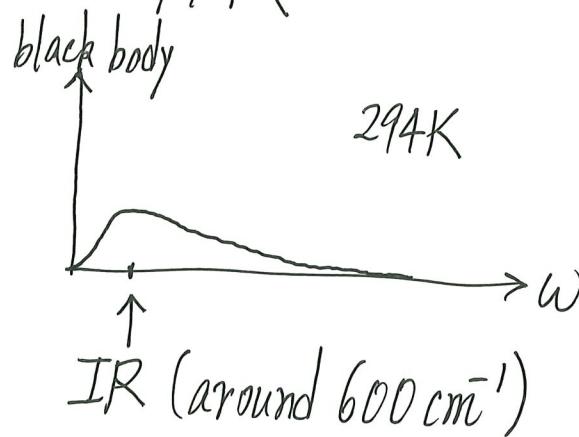
gives us  $\sim 1360 \text{ kW/m}^2$   
(solar constant)



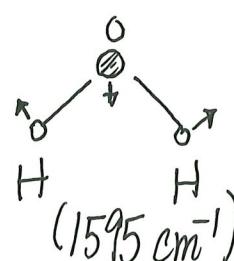
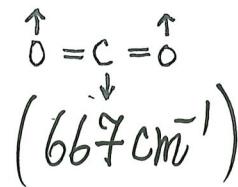
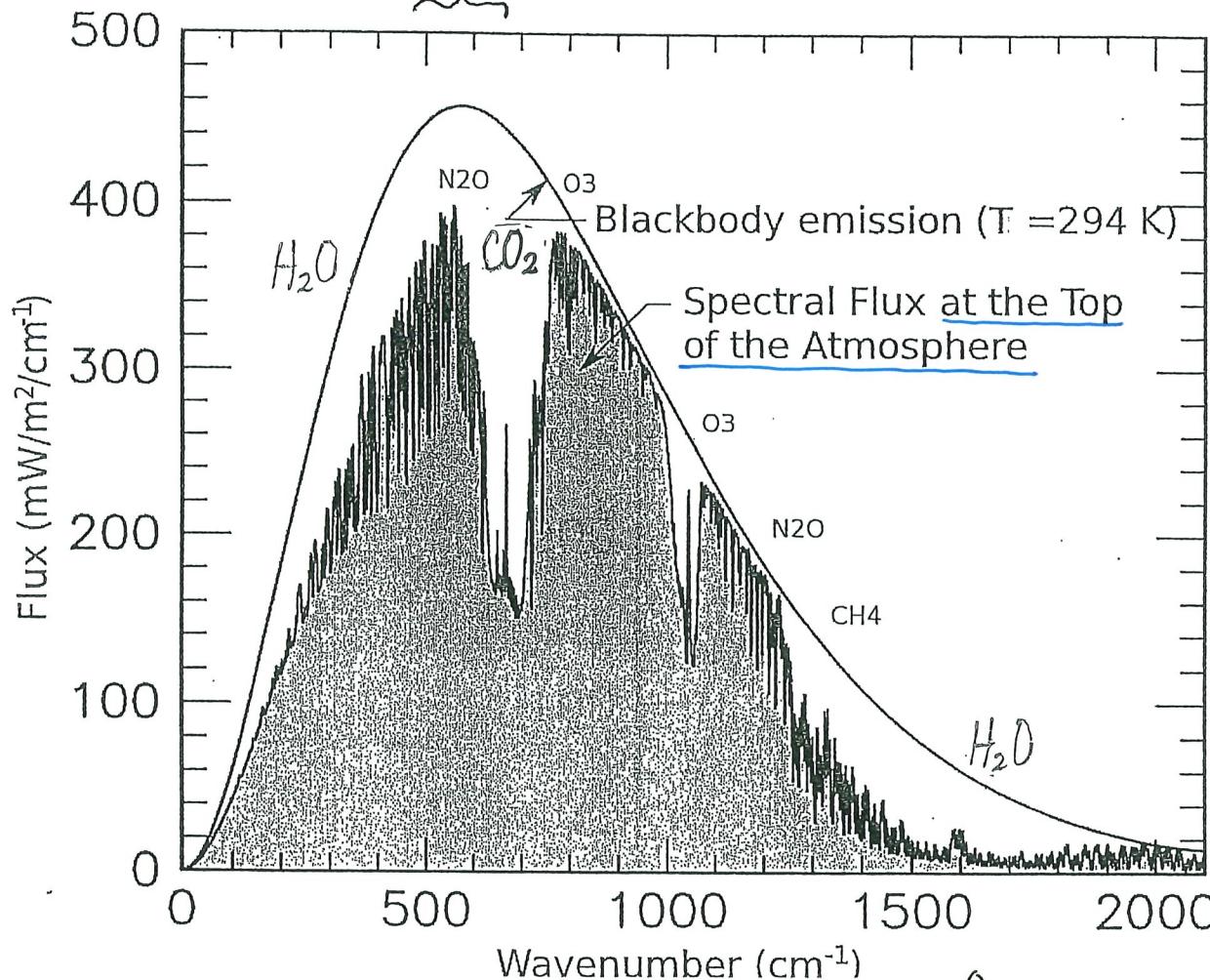
visible [our eyes evolve to fit to Sun's spectrum]

- Short wavelengths in and Long wavelengths out

Earth  $T \sim 20^\circ\text{C}$  (average)  $\sim 294\text{K}$



↗ peak ( $\sim 600 \text{ cm}^{-1}$ ) of Earth's radiations



### "Greenhouse Effect"

- $\text{CO}_2$  suppresses heat (IR) loss (keeps us warm)<sup>+</sup>
- But we don't want to be too warm

+ It is QM that keeps us warm!

- Earth's atmosphere:

~78% nitrogen, ~20% oxygen, ~0.9% argon

$\boxed{\sim 0.038\% \text{ CO}_2}$  (~2010)

Records: ~1832: 0.028%  
 ~1960: 0.032%  
 ~1970: 0.033%  
 ~1980: 0.034%  
 ~1990: 0.035%  
 ~2000: 0.037%

$\boxed{\text{Do human activities matter?}}$

[From Fayer, "Absolutely Small"]

+ How lucky we are that O<sub>2</sub> and N<sub>2</sub> are not IR active!

Ways out?

- Reduce CO<sub>2</sub> production
- Reduce CO<sub>2</sub> emission
- Let CO<sub>2</sub> be produced,  
collect & store CO<sub>2</sub>

Much research is  
on going!